

**Approximate spatially-localized HFB solutions
 for nuclei beyond the neutron drip line**

— Canonical-basis HFB method —

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Demands

1. Pairing in the **continuum** $\epsilon_F > -\frac{1}{2}\hbar\omega_{osc}$ $\frac{10^4}{2}$ nuclides
2. Long **tail** for large r $\epsilon_F > -1$ MeV halo
3. **Deformation** all but near spherical magics

Answer

Coordinate-space

Hartree-Fock-Bogoliubov

such as 3D Cartesian mesh

not HF+BCS

feasible for Skyrme force

but, How to solve ?

method	#basis	orthogonality condition	pairig force
quasi particle two basis	\propto box vol.	redundant	δ -func, + dens. dep.
canonical basis	\propto nucl. vol.	essential	+ mom. dep.

Hartree-Fock

$$|\psi\rangle = \prod_{i=1}^A a_i^\dagger |0\rangle$$

$$a_i^\dagger = \sum_s \int d^3r \psi_i(\vec{r}, s) a^\dagger(\vec{r}, s) \quad : \text{single-particle state}$$

Variation with A wavefunctions $\{\psi_i\}_{i=1, \dots, A}$
 Orthonormality condition $\langle \psi_i | \psi_j \rangle = \delta_{ij}$

$$\frac{\delta E}{\delta \psi_i^*} = h \psi_i = \epsilon_i \psi_i$$

HFB in quasi-particle method

$$|\psi\rangle = \prod_{i=1}^{\# \text{basis}} b_i |0\rangle$$

$$b_i = \sum_s \int d^3r \{ \phi_i^*(\vec{r}, s) a(\vec{r}, s) + \psi_i(\vec{r}, s) a^\dagger(\vec{r}, s) \} : \text{quasi-particle state}$$

Variation with $2 \times \# \text{basis}$ wavefunctions $\{\phi_i, \psi_i\}_{i=1, \dots, \# \text{basis}}$
 Constraints: $\langle \phi_i | \phi_j \rangle + \langle \psi_i | \psi_j \rangle = \delta_{ij}$ and $\sum \langle \psi_i | \psi_i \rangle = A$

$$\begin{pmatrix} -h & \tilde{h} \\ \tilde{h} & h \end{pmatrix} \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix} = \epsilon_i \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix}$$

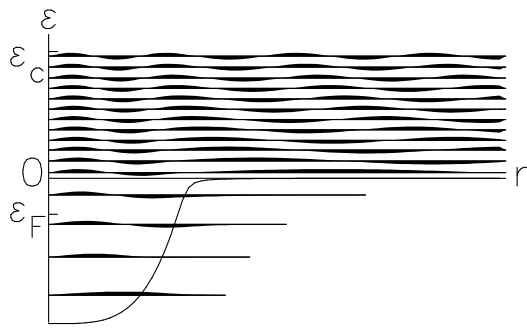
HFB in canonical-basis method

HFB solutions can be expressed in the **BCS form** (Bloch Messiah Theorem):

$$|\Psi\rangle = \prod_{i=1}^{i_{\max}} (u_i + v_i a_i^\dagger a_i^\dagger) |0\rangle$$

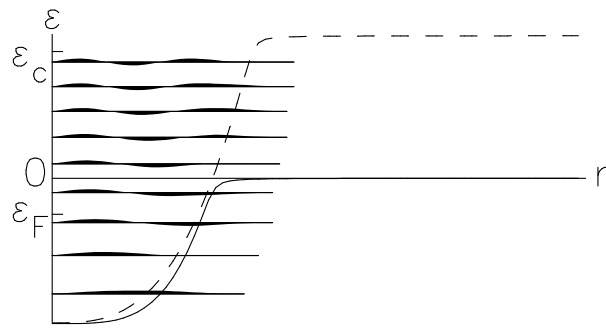
$$a_i^\dagger = \sum_s \int d^3r \psi_i(\vec{r}, s) a^\dagger(\vec{r}, s) \quad : \text{HFB canonical basis}$$

Variation with $\{\psi_i, (\bar{\psi}_i), u_i\}_{i=1, \dots, i_{\max}}$
 Constraints: $\langle \psi_i | \psi_j \rangle = \delta_{ij}$, $u_i^2 + v_i^2 = 1$ and $2 \sum v_i^2 = A$
 Exact when $i_{\max} = \frac{1}{2} \# \text{basis}$ (Bloch Messiah theorem)
 $i \leq i_{\max} = \mathcal{O}(A) \ll \# \text{basis}$: a good approximation



HARTREE-FOCK BASIS

(\approx quasi-particle states)
includes information on **excitations**.



HFB CANONICAL BASIS

One may neglect $v^2 \ll 1$ states
to describe the **ground state**.

$$\frac{\delta E}{\delta \psi_i^*} = \mathcal{H}_i \psi_i = \sum_j \lambda_{ij} \psi_j$$

Hamiltonian becomes **state-dependent** :

$$\mathcal{H}_i = v_i^2 h + u_i v_i \tilde{h}$$

h : Hartree-Fock Hamiltonian

\tilde{h} : Pairing Hamiltonian

It is \tilde{h} which governs positive-energy canonical basis.

↑
not h

↑
Resonances of h cannot replace them.

How to obtain HFB solutions in the canonical-basis form

$$|\Psi\rangle = \prod_{i=1}^K \left(u_i + v_i a_i^\dagger a_i^\dagger \right) |0\rangle \quad (1)$$

Minimize $E = \langle \Psi | H | \Psi \rangle$ by varying v_i and $\psi_i(\vec{r})$ under constraints,

$$4 \sum_{i=1}^K v_i^2 = A \quad (2)$$

$$\langle \psi_i | \psi_j \rangle = \delta_{ij} \quad (1 \leq i \leq j \leq K) \quad (3)$$

Equivalently, minimize Routhian R without constraints,

$$R = E - \epsilon_{\text{Fermi}} \cdot 4 \sum_{i=1}^K v_i^2 - \sum_{i=1}^K \sum_{j=1}^K \lambda_{ij} \{ \langle \psi_i | \psi_j \rangle - \delta_{ij} \}. \quad (4)$$

- $\lambda_{ij} = \lambda_{ji}^*$: $\frac{1}{2}K(K+1)$ constraints $\iff K^2$ multipliers
- λ_{ij} : not constants but functionals

$$\frac{\partial R}{\partial v_i} = 0 \Rightarrow v_i^2 = \frac{1}{2} - \frac{1}{2} \frac{h_{ii} - \epsilon_{\text{Fermi}}}{\sqrt{(h_{ii} - \epsilon_{\text{Fermi}})^2 + \tilde{h}_{ii}^2}}, \quad (5)$$

$$\frac{\delta R}{\delta \psi_i^*} = \mathcal{H}_i \psi_i - \sum_{j=1}^K \lambda_{ij} \psi_j - \sum_{j=1}^K \sum_{k=1}^K \frac{\delta \lambda_{jk}}{\delta \psi_i^*} \{ \langle \psi_j | \psi_k \rangle - \delta_{jk} \} = 0, \quad (6)$$

- $\mathcal{H}_i = v_i^2 h + u_i v_i \tilde{h}$ is state-dependent.
- \Rightarrow Orthogonality conditions are essential.
- \Leftrightarrow HF solution is a saddle without λ_{ij} : $h|i\rangle = \epsilon_i|i\rangle \Rightarrow \langle i|j\rangle = \delta_{ij}$

Eq. (6) at the solution (where $\langle \psi_i | \psi_j \rangle = \delta_{ij}$), \Rightarrow

$$\lambda_{ij} = \langle \psi_j | \mathcal{H}_i | \psi_i \rangle \quad \text{at the solution.} \quad (7)$$

$$\lambda_{ij} = \lambda_{ji}^* \Rightarrow \langle \psi_j | \mathcal{H}_i | \psi_i \rangle = \langle \psi_i | \mathcal{H}_j | \psi_j \rangle^* = \langle \psi_j | \mathcal{H}_j | \psi_i \rangle \quad \text{at the solution.}$$

$$\lambda_{ij} = \frac{1}{2} \langle \psi_j | (\mathcal{H}_i + \mathcal{H}_j) | \psi_i \rangle \quad : \text{ hermite everywhere} \quad (8)$$

- Acceleration of the convergence
Imaginary-time evolution method to first order in $\Delta\tau$:

$$\psi_i \rightarrow \psi_i - \Delta\tau \frac{\delta R}{\delta \psi_i^*} \quad (9)$$

More general gradient method using $\chi_i \equiv \alpha_i^{-1/2} \psi_i$:

$$\chi_i \rightarrow \chi_i - \Delta\tau \frac{\delta R}{\delta \chi_i^*} \quad (10)$$

which is equivalent to

$$\psi_i \rightarrow \psi_i - \alpha_i \Delta\tau \frac{\delta R}{\delta \psi_i^*} \quad (11)$$

We take $\alpha_i = \min(\frac{1}{v_i^2}, \frac{5}{u_i v_i})$ so that

$$\alpha_i h_i = \begin{cases} h & \text{: for deep hole states} \\ 5\tilde{h} & \text{: for high-energy particle states} \end{cases}$$

- Improved λ_{ij}

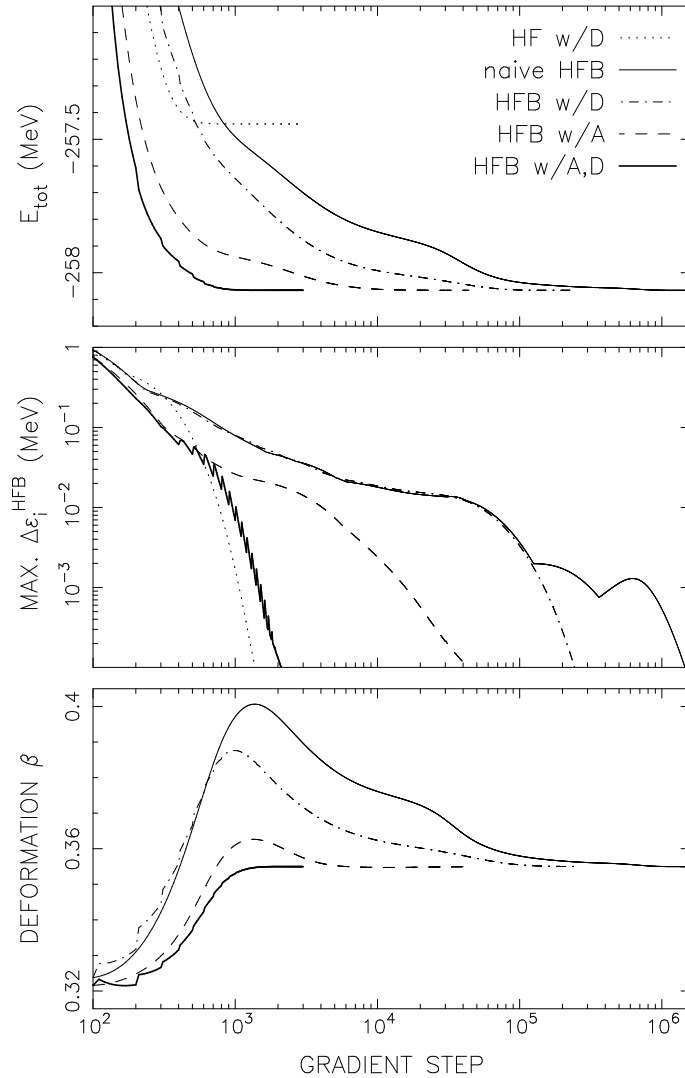
$\delta\lambda_{ij}/\delta\psi_k$ is neglected

\Rightarrow Orthogonality should be conserved

during the gradient-method iterations.

$$\lambda_{ij} = \frac{1}{\alpha_i + \alpha_j} \langle \psi_j | (\alpha_i \mathcal{H}_i + \alpha_j \mathcal{H}_j) | \psi_i \rangle$$

$$\Rightarrow \langle \psi_i | \psi_j \rangle - \delta_{ij} = \mathcal{O}((\Delta\tau)^2)$$



Skyrme force

$$\begin{aligned}\hat{v} = & t_0(1 + x_0 P_\sigma)\delta + \frac{1}{2}t_1(1 + x_1 P_\sigma)(\vec{k}^2\delta + \delta\vec{k}^2) + t_2(1 + x_2 P_\sigma)\vec{k} \cdot \delta\vec{k} \\ & + \frac{1}{6}\rho^\alpha t_3(1 + x_3 P_\sigma)\delta + iW(\vec{\sigma}_1 + \vec{\sigma}_2) \cdot \vec{k} \times \delta\vec{k}\end{aligned}$$

SIII parameters and $W = 0$ for the mean-field (HF) potential.

S=0 T=1 terms for the pairing force:

$$\hat{v}_p = v_p \frac{1 - P_\sigma}{2} \left\{ 1 - \frac{\rho}{\rho_c} - \left(\frac{\tilde{\rho}}{\tilde{\rho}_c} \right)^2 \right\} \delta - \frac{1}{2k_c^2} (\vec{k}^2\delta + \delta\vec{k}^2)$$

Hamiltonian density for even-even $N=Z$ nuclei:

$$\begin{aligned}E &= \int \mathcal{H} d\vec{r} \\ \mathcal{H} &= \frac{\hbar^2}{2m}\tau + \frac{3}{8}t_0\rho^2 + \frac{3t_1+5t_2+4t_2x_2}{16}\rho\tau + \frac{-9t_1+5t_2+4t_2x_2}{64}\rho\Delta\rho + \frac{1}{16}t_3\rho^{2+\alpha} \\ &+ \frac{1}{8}v_p \left\{ \left(1 - \frac{\rho}{\rho_c} - \left(\frac{\tilde{\rho}}{\tilde{\rho}_c} \right)^2 \right) \tilde{\rho}^2 - \frac{1}{k_c^2} (\tilde{\tau} - \Delta\tilde{\rho}) \tilde{\rho} \right\}\end{aligned}$$

where

$$\begin{aligned}\tau(\vec{r}) &= 4 \sum_{i=1}^K v_i^2 |\vec{\nabla}\psi_i(\vec{r})|^2, & \tilde{\tau}(\vec{r}) &= 4 \sum_{i=1}^K u_i v_i |\vec{\nabla}\psi_i(\vec{r})|^2, \\ \rho(\vec{r}) &= 4 \sum_{i=1}^K v_i^2 |\psi_i(\vec{r})|^2, & \tilde{\rho}(\vec{r}) &= 4 \sum_{i=1}^K u_i v_i |\psi_i(\vec{r})|^2\end{aligned}$$

Effective masses and single-particle potentials:

$$\begin{aligned}B &= \frac{\hbar^2}{2m} + \frac{3t_1+5t_2+4t_2x_2}{16}\rho, & V &= \frac{3}{4}t_0\rho + \frac{2+\alpha}{16}t_3\rho^{1+\alpha} + (\dots)\tau + (\dots)\Delta\rho - \frac{v_p}{8\rho_c}\tilde{\rho}^2 \\ \tilde{B} &= -\frac{v_p}{8k_c^2}\tilde{\rho}, & \tilde{V} &= \frac{1}{4}v_p \left\{ \left(1 - \frac{\rho}{\rho_c} - 2 \left(\frac{\tilde{\rho}}{\tilde{\rho}_c} \right)^2 \right) \tilde{\rho} - \frac{1}{2k_c^2}\tilde{\tau} + \frac{1}{k_c^2}\Delta\tilde{\rho} \right\}\end{aligned}$$

State dependent Hamiltonian:

$$\begin{aligned}h &= -\vec{\nabla} \cdot B \vec{\nabla} + V & : \text{mean-field Hamiltonian} \\ \tilde{h} &= -\vec{\nabla} \cdot \tilde{B} \vec{\nabla} + \tilde{V} & : \text{pairing Hamiltonian} \\ \mathcal{H}_i &= v_i^2 h + u_i v_i \tilde{h} & : \text{Hamiltonian of } i\text{th canonical orbital}\end{aligned}$$

t₁ term in the pairing interaction

Finite-range effects in s-wave relative motion

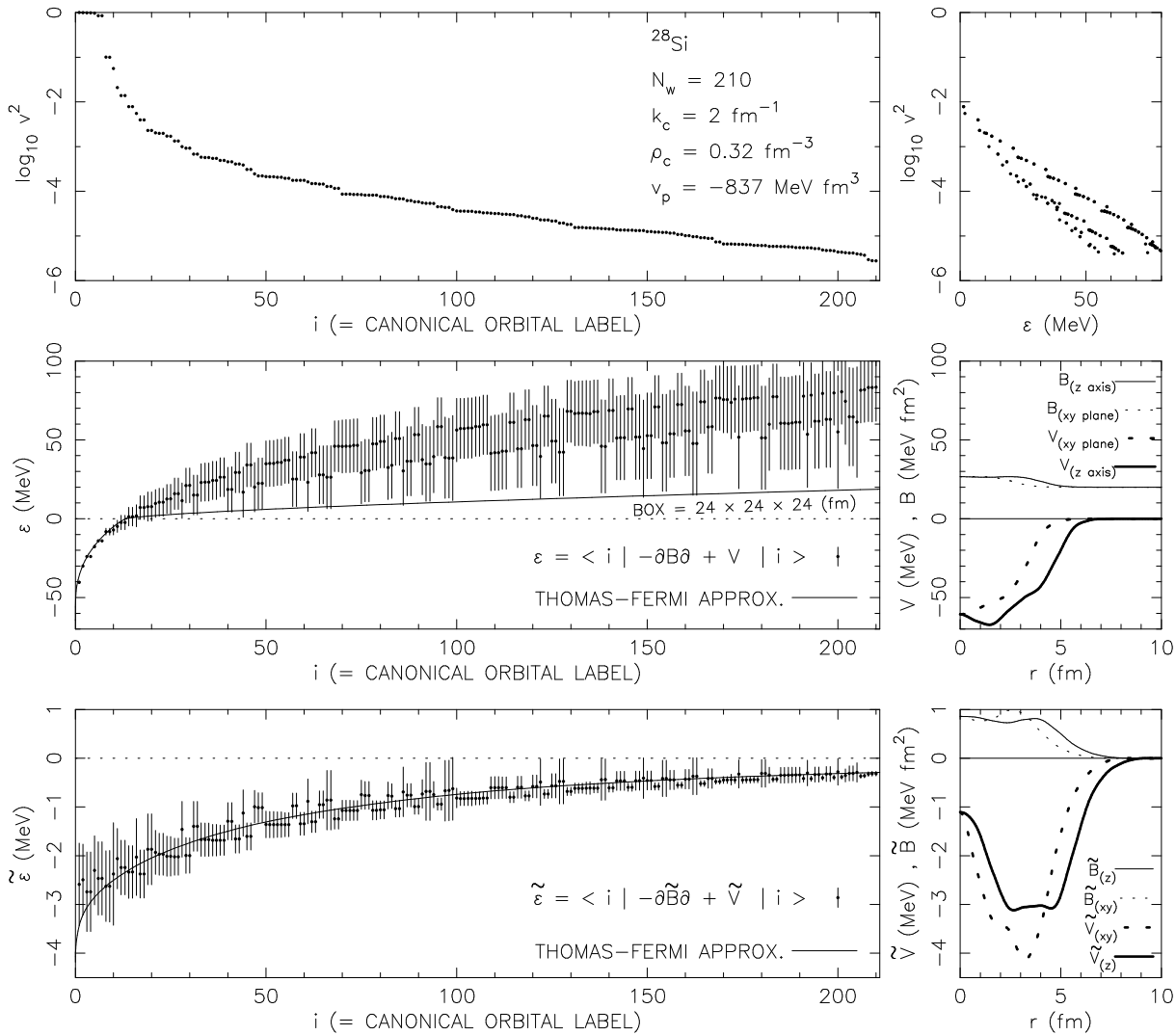


t₁ term



Kinetic energy term in \tilde{h}

1. prevents point collapse of a high-lying canonical orbital
2. allows a simple understanding of the nature of the canonical orbitals



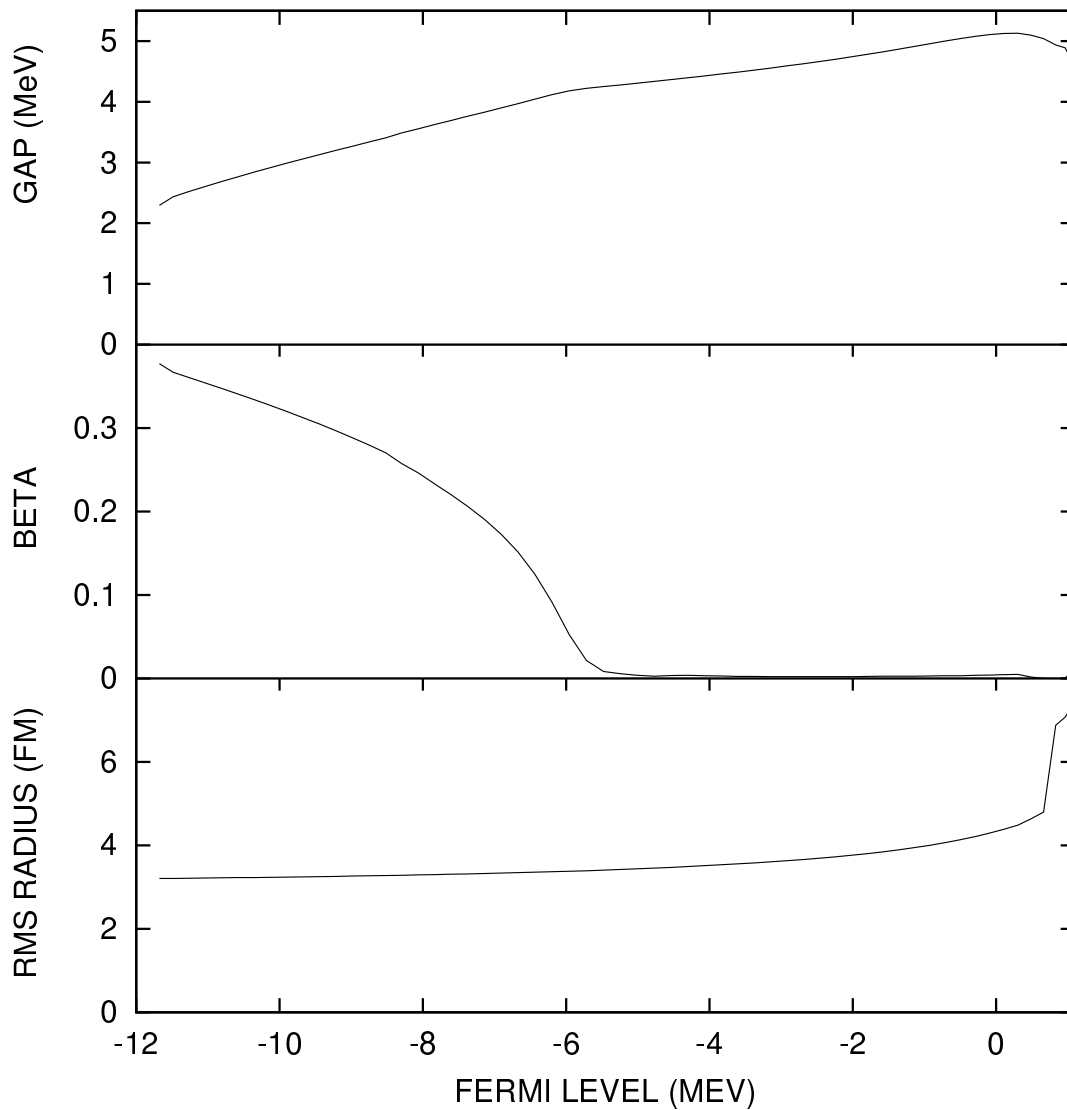
Simulation of the approach to the drip line (and beyond).

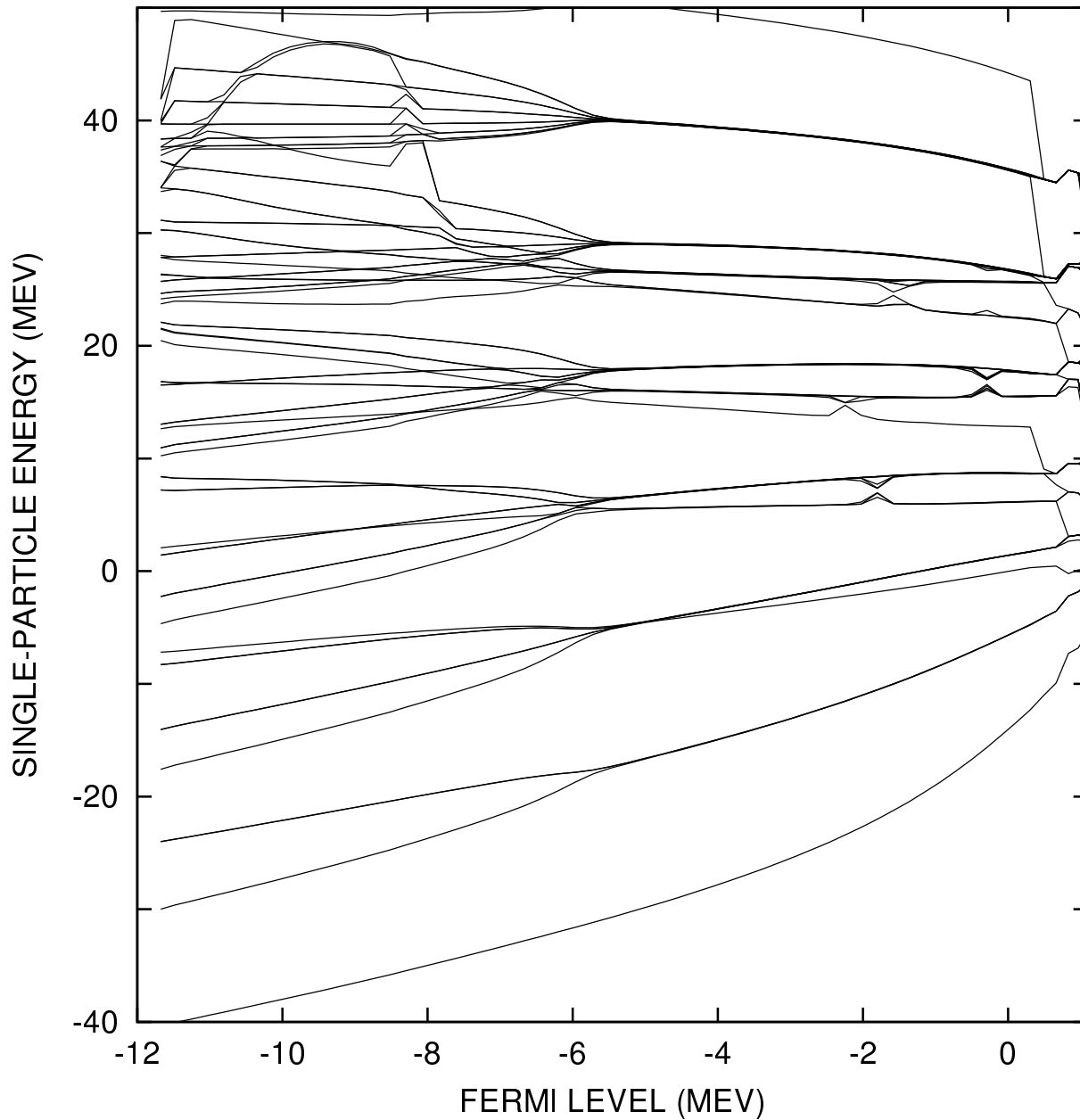
$N=Z=14$, modified SIII force ($t_0 \uparrow, t_3 \downarrow$), ls and Coulomb excluded.
 Pairing force: $v_p = -880 \text{ MeV fm}^3$, $k_c = 2 \text{ fm}^{-1}$, $\rho_c = 0.32 \text{ fm}^{-3}$.
 box: $L = 40 \text{ fm}$, $\Delta x = 0.8 \text{ fm}$, #w.f.= 70×4

$$e = \frac{\mathcal{H}}{\rho} = \frac{\hbar^2}{2m} \beta \rho^{2/3} + \frac{3}{8} t_0 \rho + \frac{1}{16} t_3 \rho^2, \quad \beta = \frac{3}{5} \left(\frac{3\pi^2}{2} \right)^{2/3}$$

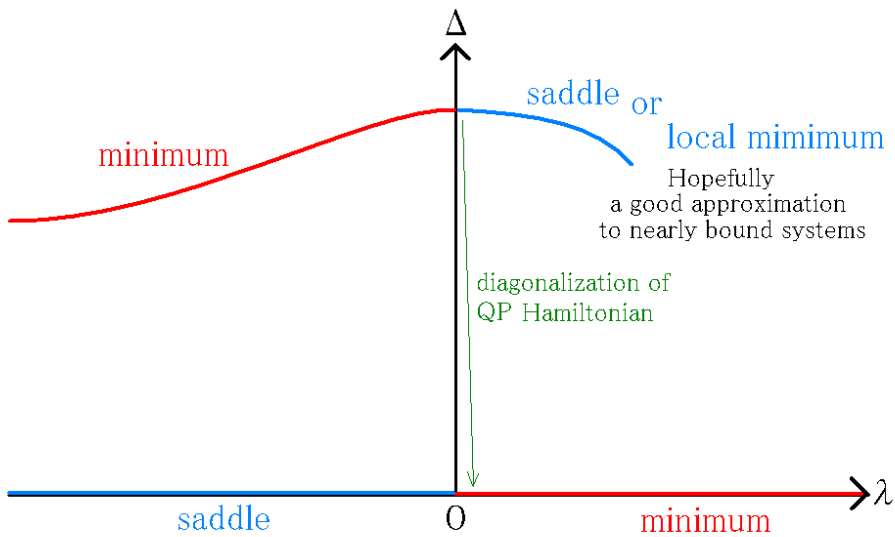
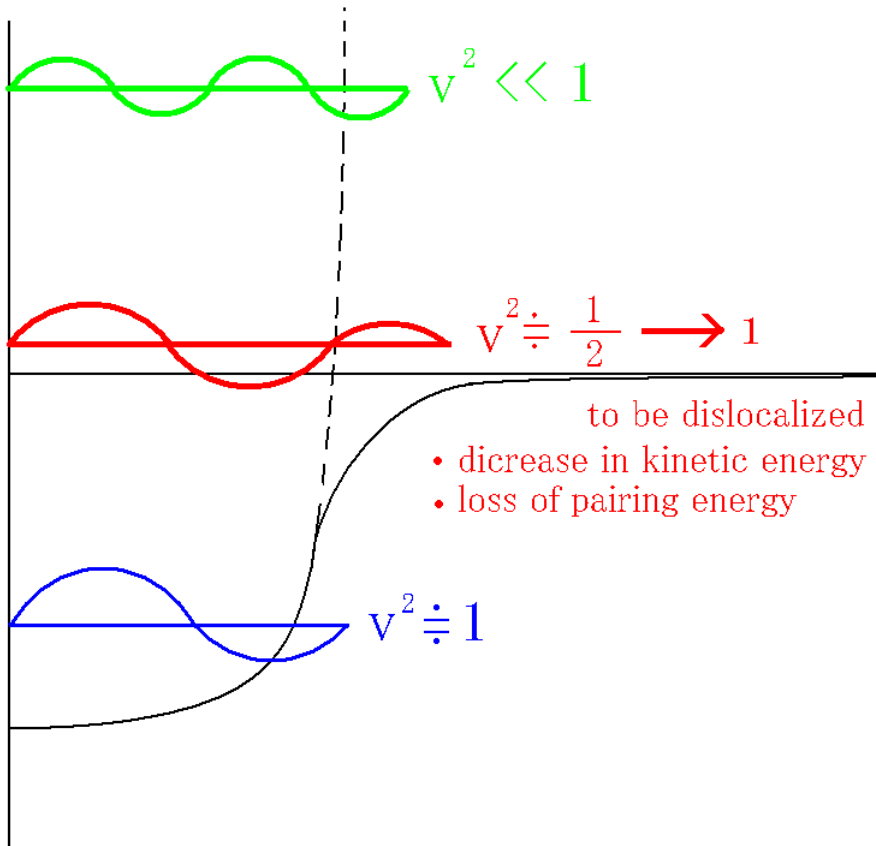
$$\frac{de}{d\rho} = 0 \text{ at } \rho = \rho_{\text{eq}}, \quad \lambda = \frac{d^2\mathcal{H}}{d\rho^2} = \frac{\mathcal{H}}{\rho} \text{ at } \rho = \rho_{\text{eq}}$$

$$t_3 = -\frac{3}{\rho_{\text{eq}}} t_0 - \frac{\hbar^2}{2m} \frac{16\beta}{3} \rho_{\text{eq}}^{-4/3}, \quad \rho_{\text{eq}} = 0.145 \text{ fm}^{-3}$$





- No drastic changes are observed in crossing the drip line.
- For $\lambda < 0$, it is related to the Pairing Anti Halo effect
- For $\lambda > 0$, the true ground state must be a dislocalized state according to quasi-particle HFB.
- Canonical-basis HFB can also give a spatially localized local-minimum solution, which may be, hopefully, a good approximation to nearly bound systems.



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